

Script

Lecture 1 — Intro

QFT

Basic object \rightarrow field $\hat{\psi}(x, p)$
 Operator-valued fn. of
 space-time, momenta..

Goal: join QM (operator, unitarity)
 with special relativity

Equations, observables are unchanged
 upon applying Lorentz transformation

L.T. def. $\lambda^{\mu\nu} \lambda^{\alpha\beta} \cdot g_{\nu\beta} = g^{\mu\alpha} = \begin{pmatrix} 1 & & & \\ & -1 & 0 & \\ 0 & & -1 & \\ & & & -1 \end{pmatrix}$
metric

Rotations in 3D + Boosts

Rotation by θ
 around z-axis

$$\begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & \cos\theta & -\sin\theta & 0 \\ 0 & \sin\theta & \cos\theta & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

Boost in z-direction
 (v in units of c)

$$\begin{pmatrix} \cosh\beta & \sinh\beta & & \\ & 1 & & \\ & & 1 & \\ \sinh\beta & & & \cosh\beta \end{pmatrix}$$

$$\cosh\beta = \frac{1}{\sqrt{1-v^2}}$$

$$\sinh\beta = \frac{\pm v}{\sqrt{1-v^2}}$$

$$\lambda : x^\mu \rightarrow x'^\mu = \lambda^{\mu\alpha} x_\alpha$$

Scalar field $\varphi(x^\mu) \rightarrow \varphi' = \varphi(\bar{\lambda}^{-1}{}^{\mu\alpha} x_\alpha)$
 (active transformation)

Vector

$$A^\mu \rightarrow \Lambda^{\mu\nu} A_\nu$$

Tensor

$$T^{\mu_1 \dots \mu_n} \rightarrow \Lambda^{\mu_1 \nu_1} \dots \Lambda^{\mu_n \nu_n} T^{\nu_1 \dots \nu_n}$$

$A_\mu B^\mu \rightarrow$ Lorentz scalar
 (paired indices stay intact;
 unpaired indices transform)

Position not a dynamical variable \rightarrow index
 of the field $\varphi(\vec{x}, t)$

Lagrangian density \rightarrow Lorentz scalar

$$KG: \mathcal{L} = \frac{1}{2} (\partial^\mu \varphi)(\partial_\mu \varphi) - \frac{1}{2} m^2 \varphi$$

$$\text{Euler-Lagrange: } \frac{\partial \mathcal{L}}{\partial \varphi} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} = 0$$

$$\hookrightarrow (\partial^2 + m^2) \varphi \Rightarrow \text{KG equation}$$

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu}, \quad F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

$$\hookrightarrow \partial_\mu F^{\mu\nu} \Rightarrow \text{Maxwell Eq. etc.}$$

Noether theorem: each continuous symmetry
 \longleftrightarrow conserved current

$$\varphi_a \rightarrow \varphi_a + \delta \varphi_a \Rightarrow \mathcal{L} \rightarrow \mathcal{L} + \delta \mathcal{L}, \quad \delta \mathcal{L} = \partial_\mu J^\mu$$

$$G \partial_\mu (\delta \varphi_a \underbrace{\frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi_a)} - J^\mu}_{\text{conserved current}}) = 0$$

*) Translation (infinitesimal)

$$x^\mu \rightarrow x^\mu - \varepsilon^\mu \Rightarrow \varphi_a(x) \rightarrow \varphi_a + \varepsilon^\nu \partial_\nu \varphi_a$$

$$\mathcal{L} \rightarrow \mathcal{L} + \varepsilon^\nu \partial_\nu \mathcal{L} \Rightarrow \underbrace{\text{Conserved current } T^{\mu\nu}}$$

Energy-momentum tensor

$$T^{\mu\nu} = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi_a)} \partial^\nu \varphi_a - g^{\mu\nu} \mathcal{L}$$

$$\partial_\mu T^{\mu\nu} = 0 \rightarrow 4 \text{ conserved comp.}$$

$$E = \int d^3x T^{00} \quad \dot{P}_i = \int d^3x T^{0i}$$

$$KG \text{ Lagrangian} \quad \mathcal{L} = \frac{1}{2} (\partial_\mu \varphi) (\partial^\mu \varphi) - \frac{1}{2} m^2 \varphi^2$$

$$\rightarrow T^{\mu\nu} = \partial^\mu \varphi \partial^\nu \varphi - g^{\mu\nu} \mathcal{L}$$

$$E = \int d^3x \left[\frac{1}{2} \dot{\varphi}^2 + \frac{1}{2} (\vec{\nabla} \varphi)^2 + \frac{1}{2} m^2 \varphi^2 \right]$$

$$\dot{P} = \int d^3x \dot{\varphi} \bar{\nabla} \varphi$$

$$\text{Inf. Lorentz transf.} \quad x^\mu_v = g^\mu_\nu + \omega^\mu_\nu$$

$$\mathcal{L} \text{ invariant} \Rightarrow J^{\mu\alpha\beta} = x^\alpha T^{\mu\beta} - x^\beta T^{\mu\alpha} \quad \underline{\omega^\mu + \omega^\nu = 0}$$

$$\text{conserved} \quad \partial_\mu J^{\mu\alpha\beta} = 0$$

(Exercise ? Find all conserved currents of K-G field)

Hamilton formalism $\Pi_a(x) = \frac{\partial \mathcal{L}}{\partial \dot{q}_a}$
conjugate mom. to q_a

→ Hamiltonian density $\mathcal{H} = \Pi_a \dot{q}_a - \mathcal{L}$

Hamiltonian $H = \int d^3x \mathcal{H}$

Real scalar: $\mathcal{L} = \frac{1}{2} \partial_{\mu}\varphi \partial^{\mu}\varphi - \frac{1}{2} V(\varphi)$

$$\mathcal{H} = \frac{1}{2} (\dot{\varphi}^2 + (\vec{\nabla}\varphi)^2 + V(\varphi))$$

→ total energy from time translation inv.

To solve it we need to diagonalize \mathcal{H}

Fourier transform $\varphi(\vec{x}, t) = \int d^3\vec{p} e^{-i\vec{p}\vec{x}} \varphi(\vec{p}, t)$

$$(\partial_t^2 - \vec{\nabla}^2 + m^2) \varphi(\vec{x}, t) = 0 \Rightarrow (\partial_t^2 + (\vec{p}^2 + m^2)) \varphi(\vec{p}, t) = 0$$

This is just H.O.

→ To quantize real scalar field: need to quantize an infinite # of H.O.

How is H.O. quantized?

QM: \hat{q} , conjugate momentum \hat{p} : $[\hat{q}, \hat{p}] = i$

$$\hat{H} = \frac{\hat{p}^2}{2m} + \frac{m\omega^2}{2} \hat{q}^2$$

Ladder op's. $\hat{a} = \sqrt{\frac{m\omega}{2}} \hat{q} + \frac{i}{\sqrt{2m\omega}} \hat{p}$

$$\hat{a}^+ = -\text{--} \quad - \quad - \text{--}$$

$$[\hat{a}^+, \hat{a}] = 1$$

$$\Rightarrow H = \omega (\hat{a}^+ \hat{a} + \frac{1}{2}) \quad [\hat{a}, H] = \omega \hat{a}$$

$$[\hat{a}^+, H] = -\omega \hat{a}^+$$

$$i \frac{d}{dt} a = [\hat{a}, H] \rightarrow a(t) = a(0) e^{-i\omega t}$$

$$\hat{a}^+(t) = \hat{a}^+(0) e^{i\omega t}$$

We can determine the spectrum

$$H|n\rangle = E_n |n\rangle \quad H|0\rangle = \frac{\omega}{2}|0\rangle \equiv E_0|0\rangle$$

n -part. state

$$(H - E_0)|n\rangle = n\omega|n\rangle$$

$$|n\rangle = \frac{1}{\sqrt{n!}} (\hat{a}^+)^n |0\rangle \quad \begin{matrix} n \text{ particles in the mode} \\ \text{with } \omega_p \end{matrix}$$

Many particles q_a, p_b $[\hat{q}_a, \hat{p}_b] = i\delta_{ab}$

Field φ_a , conj. mom π_b $[\hat{q}_a, \hat{q}_b] = [\hat{p}_a, \hat{p}_b] = 0$

$$[\varphi_a(\vec{x}, t), \pi_b(\vec{y}, t)] = i \delta^3(\vec{x} - \vec{y}) \delta_{ab} \quad \underline{\text{Equal time}}$$

$$\varphi(\vec{x}, t) = \int \frac{d^3 \vec{p}}{(2\pi)^3} \frac{1}{\sqrt{2\omega_p}} \left[\hat{a}_p e^{-ipx} + \hat{a}_p^+ e^{ipx} \right]$$

$$\dot{\pi} = \dot{\varphi} = -i \int \frac{d^3 \vec{p}}{(2\pi)^3} \sqrt{\frac{\omega_p}{2}} \left[\hat{a}_p e^{-ipx} - \hat{a}_p^+ e^{ipx} \right]$$

$$[\hat{a}_p, \hat{a}_q^+] = (2\pi)^3 \delta^3(\vec{p} - \vec{q})$$

$$\Rightarrow H = \int d^3x \mathcal{H} = \int \frac{d^3 \vec{p}}{(2\pi)^3} \omega_p \left[\hat{a}_p^+ \hat{a}_p + \frac{1}{2} (2\pi)^3 \delta(0) \right]$$

! normal ordering : all a^+ to the left
(creation)

all a to the right
(annihilation)

Why? $\rightarrow \underline{a|0\rangle = 0}$

$$H = \int \frac{d^3 \vec{p}}{(2\pi)^3} \omega_p \left[\hat{a}_p^\dagger \hat{a}_p + \frac{1}{2} (2\pi)^3 \delta^3(0) \right]$$

" " (inf. Volume)

Vacuum energy

$$H|0\rangle = \int \frac{d^3 \vec{p}}{(2\pi)^3} \omega_p \frac{1}{2} (2\pi)^3 \delta^3(0) = E_0 |0\rangle$$

" "

$$(2\pi)^3 \delta^3(0) = \lim_{L \rightarrow \infty} \left[\int_{-\frac{L}{2}}^{\frac{L}{2}} d^3 \vec{x} e^{i \vec{p} \cdot \vec{x}} \right]_{p=0} = V$$

$$\text{Energy density } \epsilon_0 = \frac{E_0}{V}$$

$$\epsilon_0 \text{ also infinite } \int \frac{d^3 \vec{p}}{(2\pi)^3} \omega_p = \infty$$

Remove vac. energy

$$\hat{H}_{KG} = \int \frac{d^3 \vec{p}}{(2\pi)^3} \omega_p \hat{a}_p^\dagger \hat{a}_p$$

EM tensor $T^{\mu\nu} = \partial^\mu \varphi \partial^\nu \varphi - g^{\mu\nu} \mathcal{L}$

$$\hat{H} = \int d^3x T^{00} \quad \hat{T}^i = \int d^3x T^{0i}$$

conserved currents (Noether theorem)

Construct Fock space

1 part. \rightarrow Hilbert space \mathcal{H}

n part. \rightarrow Fock space $\mathcal{F} = \bigoplus_n \mathcal{H}$

$$|p_1 p_2 \dots p_n\rangle = \hat{a}_{p_1}^+ \dots \hat{a}_{p_n}^+ |0\rangle$$

Properly normalized 1-p state

$$|p\rangle = \sqrt{2\omega_p} \hat{a}_p^+ |0\rangle$$

In general: to quantize real scalar field

Lagr. \rightarrow symmetries \rightarrow
conserved currents

\rightarrow identify all

\rightarrow put hats on all

\rightarrow check that all generators obey
commutation relations $[\hat{Q}_\alpha, \hat{Q}_\beta] = i f_{\alpha\beta} \hat{P}_\gamma$

\rightarrow construct Fock space

This defines unitary representation
of the Poincare group (1+transl.)

Feynman propagator

$$\Delta_F(x-y) = \langle 0 | T\{ \varphi(x), \varphi(y) \} | 0 \rangle$$

$$= \int \frac{d^4 p}{(2\pi)^4} \frac{i}{p^2 - m^2} e^{-ip(x-y)}$$

F.P. = Green's fn. of KG equation

$$(D^2 + m^2) i \Delta_F(x-y) = \delta^{(4)}(x-y)$$

Interactions $H = H_0 + \underbrace{\frac{g^{(n)}}{n!} \varphi^n}_{H_{\text{int}}} \text{ etc.}$

Perturbation theory $\hat{H} = \hat{H}_0 + \hat{H}_{\text{int}}$

t-dep. is Schröd. pict. $i \frac{d}{dt} |\psi_s\rangle = \hat{H}_s |\psi_s\rangle$

Interaction picture:

W.F. $|\psi_I(t)\rangle = e^{i\hat{H}_0 t} |\psi_s(t)\rangle$

Op. $O_I = e^{i\hat{H}_0 t} O_s e^{-i\hat{H}_0 t}$

$$i \frac{d}{dt} |\psi_I(t)\rangle = -\hat{H}_0 e^{i\hat{H}_0 t} |\psi_s\rangle + e^{i\hat{H}_0 t} i \frac{d}{dt} |\psi_s\rangle$$

$$= \left(-\cancel{\hat{H}_0} e^{i\hat{H}_0 t} + e^{i\hat{H}_0 t} (\cancel{\hat{H}_0} + \hat{H}_{\text{int}}) e^{-i\hat{H}_0 t} e^{i\hat{H}_0 t} \right) |\psi_s\rangle$$

$$= \left(\hat{H}_{\text{int}} \right)_{\perp} |\Psi_s\rangle$$

Propagator $\hat{U}(t, t_0) |\Psi_I(t_0)\rangle = |\Psi_I(t)\rangle$

$$i \frac{d}{dt} \hat{U}(t, t_0) = \hat{H}_{\text{int}} \hat{U}(t, t_0)$$

$$\hookrightarrow \hat{U}(t, t_0) = T \exp \left(-i \int_{t_0}^t \hat{H}_{\text{int}}(t') dt' \right)$$

$$T \{ O_1(t_1) O_2(t_2) \} = \begin{cases} O_1(t_1) O_2(t_2), & t_1 > t_2 \\ O_2(t_2) O_1(t_1), & t_2 > t_1 \end{cases}$$

S-matrix in- and out-states

$$\begin{cases} |i\rangle : t = -\infty \\ |f\rangle : t = +\infty \end{cases}$$

Eigenstates of free \hat{H}_0

Transition amplitude $|i\rangle \rightarrow |f\rangle$

$$S\text{-matrix} \lim_{t_{\pm} \rightarrow \pm\infty} \langle f(t_{+}) | \hat{U}(t_{+}, t_{-}) | i(t_{-}) \rangle$$

What is \hat{H}_{int} ? $\sim \phi(x_1) \phi(x_2) \dots \phi(x_n)$

$$|i\rangle = \prod_{i=1}^n (\sqrt{2\omega_{p_i}} \hat{a}_{p_i}^+) |0\rangle$$

$|f\rangle$ similar

$$H \sim a^+ a \quad \text{normal ordering} \quad \vdots \quad \vdots$$

$$+ \{ \varphi(x_1) \varphi(x_2) \dots \varphi(x_n) \}$$

$$= \overline{\varphi(x_1) \dots \varphi(x_n)} + \sum \text{all contractions}$$

$\overbrace{\varphi(x_1) \dots \varphi(x_n)}$

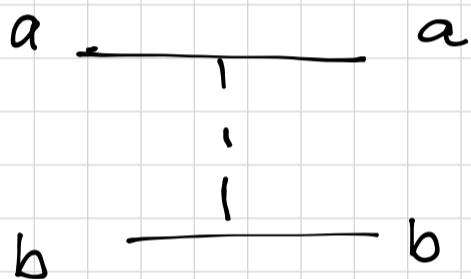
Wick's theorem

Feynman prop.

Feynman rules \rightarrow Feynman diags.

F.P. — central concept in QFT

Interaction — simple picture, eg.



contraction \rightarrow 1 p. exch.

mass of exchanged

particle \rightarrow range of int.

Unification of short and long-range forces.

Central to interaction picture: asymptotic states are eigenstates of the free Hamiltonian (\hat{H}_0). Is this still useful? We know that in QFT we cannot work with fixed # of particles. This question is answered by the LSZ reduction formula