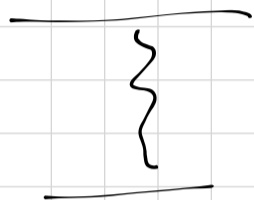


Lecture 23

We computed tree-level amplitudes for $2 \rightarrow 2$ scattering processes

Interaction at tree-level = exchange of 1 part.

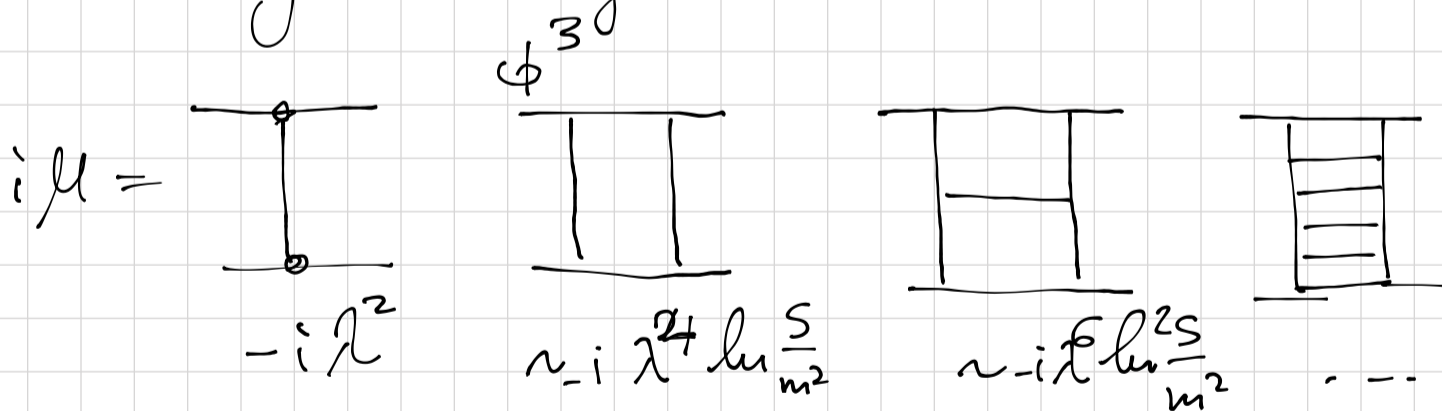


Properties of exchanged part define the interaction

($m_\gamma = 0 \rightarrow \infty$ range Coulomb force
 $m \neq 0 \rightarrow$ finite range Yukawa)

1 to 1 correspondence between the cross section and the coupling

However we have seen that whenever more particles can be produced this relation may change



$|\mathcal{M}|^2 \sim \lambda^4$

$\xrightarrow{\text{Sum over } 2+n \text{ intermediate states}}$

$\lambda^4 \left(\frac{S}{m^2}\right)^d$

The ladders was a fortunate example of finite loop diagrams

But most of the time loops = ∞

So, how can we make sense of ∞ ?

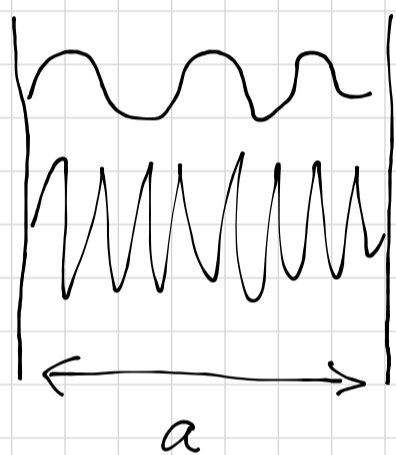
Let's take a look at ∞ that we have already seen: zero mode energy

$$H = \int \frac{d^3 \vec{k}}{(2\pi)^3} \omega_k \left[a_k^\dagger a_k + \frac{1}{2} \right]$$

$$\underline{\omega_k = |\vec{k}|}$$

$$E = \langle 0 | \int \frac{d^3 k}{(2\pi)^3} \frac{\omega_k}{2} | 0 \rangle = \infty$$

Consider a box of side "a" in the vacuum



For simplicity \rightarrow 1D

Also, assume spin-0 γ

$$E(a) \rightarrow F(a) = - \frac{dE}{da}$$

"Standing waves" $\omega_n = \frac{\pi}{a} n$

$$E(a) = \frac{1}{2} \sum_n \omega_n = \frac{\pi}{2a} \sum_n n = \infty$$

Ramanujan summation

$$\sum_n n = -\frac{1}{12}$$

$$A = 1 - 1 + 1 - 1 + \dots = \frac{1}{2}$$

$$\rightarrow 1 - A = 1 - (1 - 1 + 1 - 1 \dots) = A \Rightarrow A = \frac{1}{2}$$

$$B = 1 - 2 + 3 - 4 + 5 - 6 + \dots = \frac{1}{4}$$

$$\rightarrow A - B = \frac{1 - 1 + 1 - 1 + \dots}{(1 - 2 + 3 - 4 + \dots)} = 1 - 2 + 3 - 4 + \dots = B \Rightarrow B = \frac{1}{2}A = \frac{1}{4}$$

$$C = 1 + 2 + 3 + 4 + 5 + \dots = -\frac{1}{12}$$

$$C - B = \cancel{1} + 2 + \cancel{3} + 4 + \cancel{5} + 6 + \dots - (\cancel{1} - 2 + \cancel{3} - 4 + \cancel{5} - 6 + \dots)$$

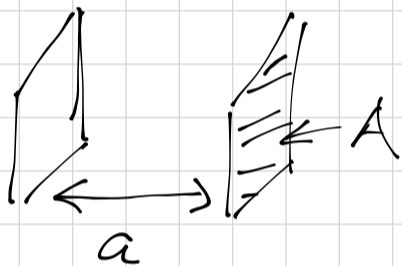
$$= 4 + 8 + 12 + \dots = 4(1 + 2 + 3 + \dots) = 4C$$

$$C = -\frac{1}{3}B = -\frac{1}{12}$$

$$E(a) = -\frac{\pi}{24a} \longrightarrow F(a) = -\frac{\pi}{24a^2}$$

3D + 2 pol. of the γ

$$F(a) = -\frac{\pi \hbar c}{240a^4} A$$



$$E(a) = \frac{1}{2} \sum_n \omega_n(a)$$

Add a regulator $f(n)$: suppress higher modes

$$E(a) = \frac{1}{2} \sum_n \left(\frac{\omega_n}{\mu}\right)^{-s} \cdot \omega_n = \frac{1}{2} \left(\frac{\pi}{a}\right)^{1-s} \mu^s \underbrace{\sum_n \frac{1}{n^{s-1}}}_{\text{Riemann zeta-fn.}}$$

Riemann zeta-fn. $\zeta(s-1)$

If $s=0 \longrightarrow$ we recover the original sum $\sum_n n$

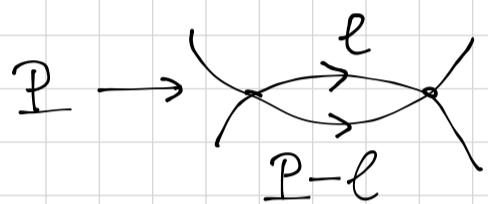
However, $\zeta(-1) = -\frac{1}{12}$

\hookrightarrow take limit $s \rightarrow 0$ $E(a) = -\frac{\pi}{24a}$

Chapter 15 of Schwartz "Casimir effect"

For any reg. fn. $f(n)$ the result for the Casimir force is the same!

We got a finite number by subtracting ∞ .
1-loop in ϕ^4



$$i\mu^{(2)} = \frac{\lambda^2}{2} \int \frac{d^4 l}{(2\pi)^4} \frac{1}{l^2 - m^2} \frac{1}{(P-l)^2 - m^2}$$

$$2 \operatorname{Im} \mu^{(2)} = \frac{\lambda^2}{2} \int \frac{d^4 l}{(2\pi)^4} 2\pi \delta(l^2 - m^2) 2\pi \delta((P-l)^2 - m^2)$$

$$P^2 = s$$

$$P^0 = \sqrt{s}$$

$$P = 0$$

$$\Theta(l^0) \Theta(P^0 - l^0)$$

$$\Rightarrow \operatorname{Im} \mu^{(2)} = \frac{\lambda^2}{32\pi} \sqrt{1 - \frac{4m^2}{s}} \Theta(s - 4m^2)$$

\int_{DR}

$$\mu^{(2)} = \frac{1}{\pi} \int_{4m^2}^{\infty} \frac{ds'}{s' - s - i\epsilon} \operatorname{Im} \mu^{(2)}(s')$$

$$\mu^{(2)}(\Lambda) = \frac{1}{\pi} \int_{4m^2}^{\Lambda^2} \frac{ds'}{s' - s - i\epsilon} \operatorname{Im} \mu^{(2)}(s')$$

$$= + \frac{\lambda^2}{32\pi^2} \ln \left(\frac{\Lambda^2}{s} \right)$$

$\Lambda \rightarrow \infty \Rightarrow \mu^{(2)}(\Lambda)$ diverges log.

$$\mu^{(2)}(s_1) - \mu^{(2)}(s_2) = \frac{\lambda^2}{32\pi^2} \ln \left(\frac{s_2}{s_1} \right)$$

$$\frac{dG}{d\Omega} \sim |\mu^{(n+2)}|^2 \sim \underline{f(\lambda)}$$

$$A\left(\frac{\lambda}{f_a}\right) \quad |A| \leq A^{\text{exp.}}$$

$$\Leftrightarrow \left|\frac{\lambda}{f_a}\right| \leq \dots$$

We measure

$$i\mu = -i\lambda - \frac{i\lambda^2}{32\pi^2} \ln S + C_\infty + \dots$$

$$C_\infty = \frac{i\lambda^2}{32\pi^2} \ln \Lambda^2$$

Define renormalized coupling λ_R

$$\underline{\lambda_R} \equiv -\mu(s_0) = \lambda - \frac{\lambda^2}{32\pi^2} \ln \frac{s_0}{\Lambda^2} + \dots$$

$$\lambda = \lambda_R + a \lambda_R^2 + \dots$$

$$\lambda_R = (\lambda_R + a \lambda_R^2 + \dots) - \frac{(\lambda_R + a \lambda_R^2 + \dots)^2}{32\pi^2} \ln \frac{s_0}{\Lambda^2} + \dots$$

$$\rightarrow a = -\frac{1}{32\pi^2} \ln \frac{s_0}{\Lambda^2}$$

$$\lambda = \lambda_R - \frac{\lambda_R^2}{32\pi^2} \ln \frac{s_0}{\Lambda^2}$$

\Downarrow

$$\mu(s) = -\lambda - \frac{\lambda^2}{32\pi^2} \frac{s}{\Lambda^2} + \dots$$

$$= -\lambda_R + \frac{\lambda_R^2}{32\pi^2} \ln \frac{s_0}{\Lambda^2} + \dots$$

$$\begin{aligned}
 & - \frac{\lambda_R^2}{32\pi^2} \ln \frac{s}{\Lambda^2} + \dots \\
 & = -\lambda_R - \frac{\lambda_R^2}{32\pi^2} \ln \frac{s}{s_0} + \dots
 \end{aligned}$$

$$\underline{\lambda_R = -\mu(s_0)}$$

Again: what we did is subtracting the ∞

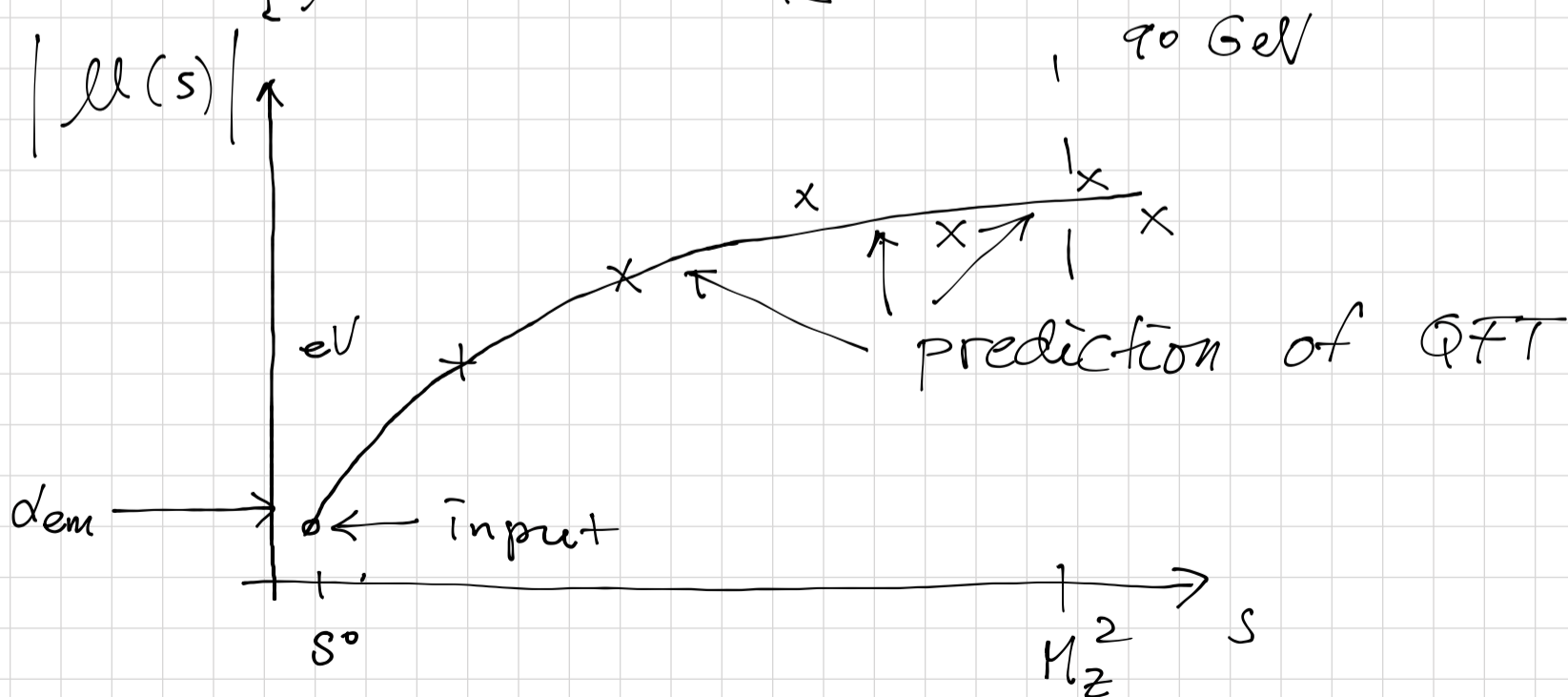
Counterterm

$$\mathcal{L} = -\frac{1}{2} \phi \square \phi - \frac{\lambda_R}{4!} \phi^4 - \frac{\delta\lambda}{4!} \phi^4$$


$$\mu(s) = -\lambda_R - \delta\lambda - \frac{\lambda_R^2}{32\pi^2} \ln \frac{s}{\Lambda^2} + \dots$$

$$\delta\lambda = -\frac{\lambda_R^2}{32\pi^2} \ln \frac{s_0}{\Lambda^2}$$

$$\begin{aligned}
 \hookrightarrow \mu(s) &= -\lambda_R - \frac{\lambda_R^2}{32\pi^2} \ln \frac{s}{s_0} + \dots \\
 \mu(s_0) &= -\lambda_R
 \end{aligned}$$

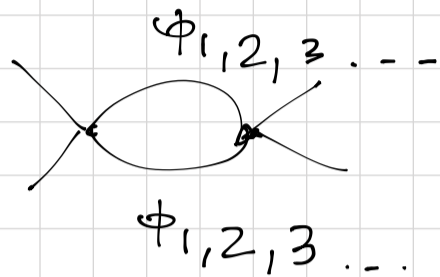


Measuring this curve \rightarrow substance of precision tests of Standard Model

In the example 

The result depends on λ
on the spectrum

If different types of part.



If you find a deviation $\mu(s)^{\text{exp}} \neq \mu(s)^{\text{th.}}$

Possibility: the spectrum $\phi_{1,2,3} \dots$ incomplete

\hookrightarrow add a new particle ϕ_{n+1} :

$$\mu^{\text{exp}} = \mu^{\text{th}}$$

Example: t-quark $m_t \cong 170 \text{ GeV}$

Predicted from a comparison of PT to

exp. corrections $\sim \ln\left(\frac{m_t^2}{s}\right)$

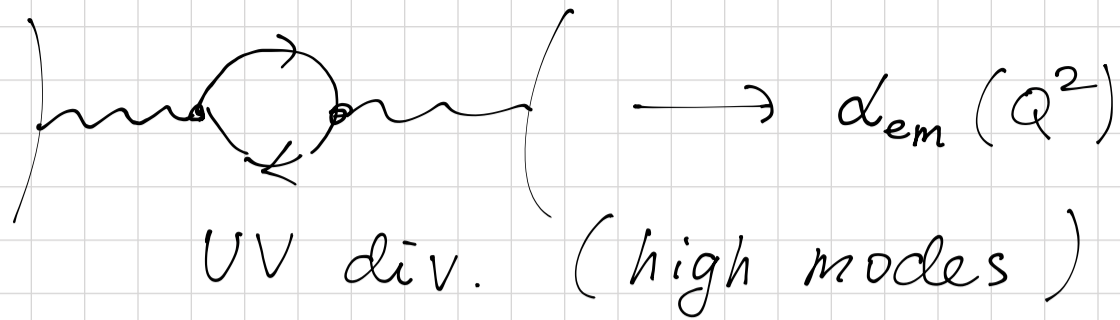
Example 2: Higgs boson $m_H \cong 125 \text{ GeV}$
LHC 2012

\hookrightarrow corrections tell you how obs.
depend on masses

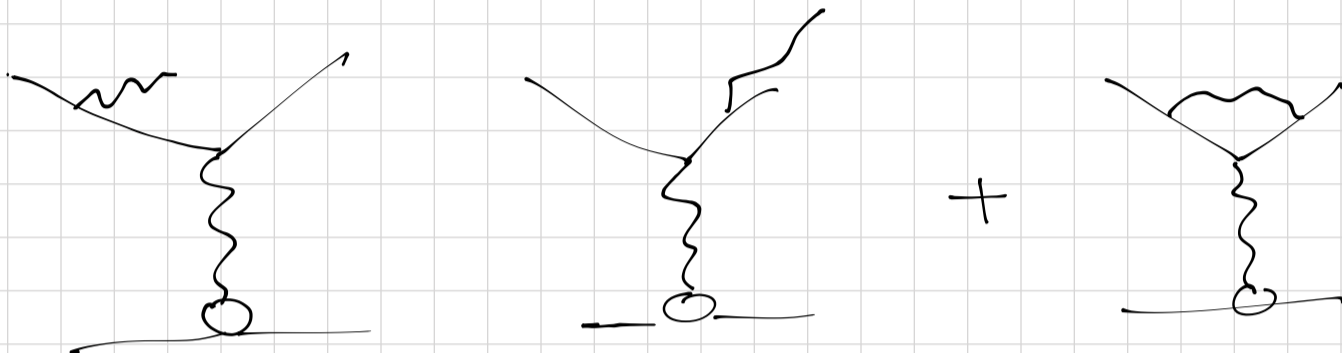
\hookrightarrow global fits \rightarrow determine
parameters

Higgs mass had to be in a
rel. narrow window around 125
GeV

→ Vacuum polarization



→ IR div. → γ 's are massless



$$\mathcal{M}^{(0)}(e+p \rightarrow e+p)$$

$$= -\frac{e^2}{t} \bar{u}(k') \gamma^\mu u(k) \cdot J_\mu$$



Consider $\mathcal{M}^{(1)}(e+p \rightarrow e+p+\gamma)$

↳ carries away energy

$$\mathcal{M}^{(1)} = -\frac{e^2}{t} \bar{u}(k') \frac{\not{\epsilon}^*(q_\gamma) (\not{k}' - \not{q}_\gamma + m_e) \gamma^\mu}{(k' - q_\gamma)^2 - m_e^2} u(k) \cdot J_\mu$$

Study the soft limit: $q_\gamma \rightarrow 0$

$$\underbrace{(k' - q_\gamma)^2 - m_e^2}_{\checkmark} = -2k'q_\gamma \sim \omega_\gamma$$