

Lecture 5

1

We introduced propagator

$$D(x-y) = \langle 0 | \phi(x) \phi(y) | 0 \rangle$$
$$= \int \frac{d^3 \vec{p}}{(2\pi)^3} \frac{e^{-i p(x-y)}}{2 \sqrt{\vec{p}^2 + \omega^2}}$$

Causality: for any $(x-y)^2 = (x^0-y^0)^2 - (\vec{x}-\vec{y})^2 < 0$

$$\Delta(x-y) = D(x-y) - D(y-x) = 0$$

Introduced Feynman propagator

$$\Delta_F = \begin{cases} D(x-y), & x^0 > y^0 \\ D(y-x), & y^0 > x^0 \end{cases}$$

We will see that Δ_F → central role in QFT interaction

Now $D(x-y) - D(y-x) = 0$ for $(x-y)^2 < 0$

$$D(x-y) \neq 0$$

choose $x^0 = y^0$

What is the meaning

$$D(x-y) = \int \frac{d^3 \vec{p}}{(2\pi)^3} \frac{e^{+i \vec{p}(\vec{x}-\vec{y})}}{2 \sqrt{\vec{p}^2 + \omega^2}}$$

$$|\vec{x}-\vec{y}| = r$$

$$p_z \parallel \vec{x}-\vec{y}$$

$$= \int_0^\infty \frac{p^2 dp}{(2\pi)^3} \int_0^{2\pi} d\varphi \int_{-1}^1 d\cos\theta \frac{e^{i p r \cos\theta}}{2 \sqrt{p^2 + \omega^2}}$$

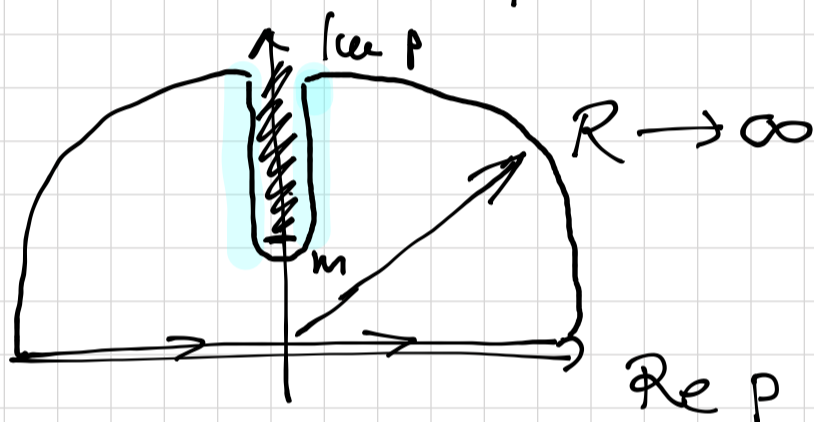
$$= \frac{-i}{8\pi^2 r} \int_0^{\infty} \frac{p dp}{\sqrt{p^2 + \omega^2}} (e^{ipr} - e^{-ipr})$$

$\int \rightarrow$ Complex plane

$e^{ipr} \rightarrow$ upper plane ($\text{Im } p > 0, e^{-\text{Im } p r}$)
 $e^{-ipr} \rightarrow$ lower $\text{Im } p < 0$

Attn: for $p = i\omega, |p| > \omega$
 $p^2 + \omega^2 < 0$
 $\sqrt{p^2 + \omega^2} = \pm i \sqrt{|p^2 + \omega^2|}$

1st



$$\oint = \int_0^{\infty} \dots + \int_R \dots + \int_{\text{cut}}$$

$$\int_R: p = R \cdot e^{i\varphi} \Rightarrow \int_R = \int_0^{\pi/2} \frac{R^2 e^{2i\varphi} dy}{\sqrt{R^2 e^{2i\varphi} + \omega^2}}$$

$e^{iRr \cos \varphi} \quad e^{-Rr \sin \varphi}$
 $R \rightarrow \infty \rightarrow 0$

$$\int_{\text{cut}}: = \int_{i\infty + \epsilon}^{i\omega + \epsilon} + \int_{i\omega - \epsilon}^{i\infty - \epsilon}$$

$p \rightarrow iy$

$$\sqrt{(i|p| \pm \epsilon)^2 + \omega^2} = \pm i \sqrt{|p|^2 - \omega^2}$$

$$\int_{\text{cut}} = \frac{2}{8\pi^2 r} \int_m^\infty \frac{y dy}{\sqrt{y^2 - u^2}} e^{-yr} \quad (\times 2)$$

$\oint = 0$: no singularities inside

$$D(r) \sim \frac{1}{r} \int_1^\infty \frac{x dx}{\sqrt{x^2 - 1}} e^{-mr x} \quad y = m \cdot x \quad K_1(mr)$$

Asymptotic behavior $D(r) \quad r \rightarrow \infty$

$$r \rightarrow \infty \quad D(r) \sim \frac{e^{-mr}}{r} \rightarrow \text{Yukawa p. (short-range)}$$

$$m = 0 \quad D(r) \sim \frac{1}{r} \rightarrow \text{Electrostatic Gravity}$$

Field leaks outside the lightcone

Classical p. $\sim \frac{1}{r} \rightarrow$ massless field

Short-range (strong, weak) $\sim \frac{e^{-mr}}{r}$ massive field

More to this \rightarrow next topic

Before going there : non relativistic limit

$$K-G \text{ eq. } (\partial_\mu \partial^\mu + m^2) \psi = 0$$

$$\psi(\vec{x}, t) = \underline{e^{-imt}} \Psi(\vec{x}, t)$$

$$\ddot{\psi} - \vec{\nabla}^2 \psi + m^2 \psi = \cancel{\psi} - 2im \dot{\psi} - \vec{\nabla}^2 \psi = 0$$

Non-rel. limit $|\vec{p}| \ll m$

$$\hookrightarrow |\dot{\psi}| \ll m |\psi|$$

slowly-varying field

$$i \frac{\partial \psi}{\partial t} = - \frac{\vec{\nabla}^2}{2m} \psi$$

$$\mathcal{L} = i \psi^* \dot{\psi} - \frac{1}{2m} \overline{\nabla \psi^*} \nabla \psi$$

Symmetry $\psi \rightarrow e^{i\alpha} \psi$ $\psi \rightarrow \psi + i\alpha \psi$
 $\psi^* \rightarrow \psi^* - i\alpha \psi^*$

\hookrightarrow Conserved Noether current

$$Q = \int d^3 \vec{x} j^0 \quad j^\mu = \left(-\psi^* \psi, i \frac{1}{2m} (\psi^* \nabla \psi - \psi \nabla \psi^*) \right)$$

$$\rightarrow Q = \int d^3 \vec{x} : \psi^* \psi : = \text{constant}$$

\hookrightarrow particle number

- conserved # of particles
- complex field but only 1 kind (no antiparticles)

$$\psi^*(\vec{x}) = \int \frac{d^3 \vec{p}}{(2\pi)^3} a_p^\dagger e^{-i\vec{p}\vec{x}}$$

$$|\vec{x}\rangle = \psi^*(\vec{x}) |0\rangle$$

$$\hat{X}^i = \int d^3 \vec{x} \vec{x} \psi^* \psi$$

Superposition of 1-particle states $|\psi\rangle = \int d^3 \vec{x} \psi(\vec{x}) |\vec{x}\rangle$

$$\hat{P} |\psi\rangle = \int d^3\vec{x} \left(-i \frac{\partial \psi}{\partial x^i} \right) |\vec{x}\rangle$$

$$i \frac{\partial \psi}{\partial t} = -\frac{\vec{\nabla}^2}{2m} \psi$$

$$Q = \int d^3\vec{x} |\psi(\vec{x})|^2$$

Schr. WF with
statistical
interpretation

Interacting fields

We diagonalized the Hamiltonian
in normal modes $\hat{a}_p, \hat{a}_p^\dagger$;

Every mode evolves independently

↳ free theory $\rightarrow \mathcal{L}$ quadratic in ϕ

\Rightarrow Interaction = higher orders in ϕ

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi - \sum_{n \geq 3} \frac{\lambda_n}{n!} \phi^n$$

λ_n coupling constants

$$[S] = \hbar = 0 \quad (\text{means } E^0)$$

$$S = \int d^4x \mathcal{L} \Rightarrow [\mathcal{L}] = 4$$

$$[\partial_\mu] = 1 \quad [\phi] = 1$$

$$\Rightarrow [\lambda_n] = 4 - n$$

$[\lambda_3] = 1 \Rightarrow$ "relevant interaction"
 High energies: $\frac{\lambda_3}{E} \ll$
 large at low E

$[\lambda_4] = 0 \Rightarrow$ "marginal"
 $|\lambda_4| \ll 1$
 \rightarrow small effects

$[\lambda_{n \geq 5}] = n - 4 \Rightarrow$ "irrelevant"
 $\lambda_n \sim \frac{E^{n-4}}{\Lambda^{n-4}}$
 low energies \ll
 blows up at large E

The power of dimensional analysis

Imagine we extend our R-G theory to
 very high scale Λ $[\Lambda] = 1$
 What will it look at $E \ll \Lambda$?

↓

$$\frac{\mathcal{L}_0}{\Lambda^4} = \sum_{n \geq 5} \frac{1}{n!} g_n \frac{E^{n-4}}{\Lambda^{n-4}} \dots$$

$\Lambda \rightarrow \infty$
 interaction $\rightarrow 0$

Weak interactions

\hookrightarrow Fermi int.

$$\frac{G_F}{\sqrt{2}} \bar{\psi} \gamma^\mu (1 - \gamma_5) \psi \quad \bar{\psi} \gamma_\mu (1 - \gamma_5) \psi$$

$$[G_F] = -2 \left(\frac{1}{\Lambda^2} \right)$$

$$G_F \sim 1 \cdot 10^{-5} \text{ GeV}^{-2}$$

↳ discovery of W^\pm with $m_W \sim 80 \text{ GeV}$

$$G_F \sim \frac{4\pi d}{M_Z^2} \quad Z^0 \text{ with } m_Z \sim 90 \text{ GeV}$$

How to solve QFT w. interactions?

1. Perturbation theory

↳ Feynman diagrams infinite series

2. Variational methods

3. Monte Carlo simulations (eg. on discretized space-time → lattice)

A. Exact solutions

1+1 dim (Bethe ansatz)

2+1 dim Topological QFT

Large N_c limit

Perturbation theory

$$\hat{H} = \hat{H}_0 + \hat{H}_{int}$$

$$H = T + V$$

$$L = T - V$$

+ - dep. Schrodinger eq. $i \frac{d}{dt} |\psi_s\rangle = \hat{H} |\psi_s\rangle$

Interaction picture

All states and operators are transformed as

$$|\Psi_I(t)\rangle = e^{i\hat{H}_0 t} |\Psi_S(t)\rangle$$

$$O_I = e^{i\hat{H}_0 t} O_S e^{-i\hat{H}_0 t}$$

$$i \frac{d}{dt} |\Psi_I(t)\rangle = -\hat{H}_0 e^{i\hat{H}_0 t} |\Psi_S\rangle + e^{i\hat{H}_0 t} \underbrace{i \frac{d}{dt} |\Psi_S\rangle}_{\hat{H} |\Psi_S\rangle}$$

$$= \left(\cancel{-\hat{H}_0 e^{i\hat{H}_0 t}} + e^{i\hat{H}_0 t} \left(\cancel{\hat{H}_0} + \hat{H}_{int} \right) e^{-i\hat{H}_0 t} e^{i\hat{H}_0 t} \right) |\Psi_S\rangle$$

$\hat{H}_{int} \big|_I(t)$

$$i \frac{d}{dt} |\Psi_I(t)\rangle = \hat{H}_{int}(t) |\Psi_I(t)\rangle$$

Propagator $\hat{U}(t, t_0): |\Psi_I(t)\rangle = \hat{U}(t, t_0) |\Psi_I(t_0)\rangle$

$$i \frac{d}{dt} \hat{U}(t, t_0) = \hat{H}_{int}(t) \hat{U}(t, t_0)$$

$$U(t, t_0) \neq e^{-i \int_{t_0}^t \hat{H}_{int}(t') dt'}$$

$$\text{Exp}[Op.] = \mathbb{1} + Op. + \frac{1}{2}(Op)^2 + \dots$$

$$\frac{(-i)^2}{2} \left(\int_{t_0}^t \hat{H}_{int}(t') dt' \right)^2$$

Back into diff. eq./

$$\frac{d\hat{U}}{dt} = \dots - \frac{1}{2} \left[\int_{t_0}^t H_{int}(t') dt' \right] H_{int}(t)$$

$$- \frac{1}{2} H_{int}(t) \left[\dots \right] + \dots$$

1st term not ok.

$$[H(t), H(t')] \neq 0 \quad t \neq t'$$

Correct solution:

$$\hat{U}(t, t_0) = T \exp \left(-i \int_{t_0}^t H_{int}(t') dt' \right)$$

$$T(O_1(t_1) O_2(t_2)) = \begin{cases} O_1(t_1) O_2(t_2), & t_1 > t_2 \\ O_2(t_2) O_1(t_1), & t_2 > t_1 \end{cases}$$

$$i \frac{\partial}{\partial t} T \left(\exp \left(-i \int_{t_0}^t H_I(t') dt' \right) \right)$$

$$= T \left[\underbrace{H_I(t)}_{\text{circled}}, \exp \left(-i \int_{t_0}^t H_I(t') dt' \right) \right]$$

Trivial proof

Dyson formula

