

## Cosmology and General Relativity: HW 4

Turn in to “Wittig” (Hausaufgaben) mailbox in KPH by **noon, 13 June 2019**

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### Problem 1 Commutators and Jacobi identity:

- a. Show that equipping the space of vector fields on manifold  $M$  with the commutator  $[X, Y] = XY - YX$  forms an algebra by showing that the commutator is bilinear and closed.<sup>1</sup> 2pt.
- b. For vector fields  $X, Y$ , and  $Z$  (all smooth) on manifold  $M$ . Verify the Jacobi identity: 1pt.

$$[[X, Y], Z] + [[Y, Z], X] + [[Z, X], Y] = 0 .$$

- c. Let  $V_1, \dots, V_n$  be a set of vector fields on the  $n$ -dimensional manifold  $M$  such that they form a basis of the tangent space at all points on  $M$ . We can define the *structure constants* with respect to this basis

$$[V_i, V_j] \equiv f_{ij}{}^k V_k$$

with the symmetry condition  $f_{ij}{}^k = -f_{ji}{}^k$ . Express the Jacobi identity in terms of the structure constants. 1pt.

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### Problem 2 Properties of the covariant derivative:

A covariant derivative (also called an *affine connection*)  $\nabla$  is defined as a map from two vector fields to one vector field,

$$\begin{aligned} \nabla : \text{Vect}(M) \times \text{Vect}(M) &\rightarrow \text{Vect}(M) \\ (X, Y) &\mapsto \nabla_X Y, \end{aligned}$$

which satisfies the following rules,

$$\begin{aligned} \nabla_X(Y + Z) &= \nabla_X Y + \nabla_X Z, & \nabla_{(fX)} Y &= f \nabla_X Y, \\ \nabla_{(X+Y)} Z &= \nabla_X Z + \nabla_Y Z, & \nabla_X(fY) &= X[f]Y + f \nabla_X Y \text{ (Leibniz)}, \end{aligned}$$

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<sup>1</sup>Algebraic aside: the tangent space at a point  $p \in M$  forms a *vector space* over the field  $\mathbb{R}$ , and equipping a vector space with a bilinear operator defines an *algebra* (over a field). However, the space of vector fields actually forms a *module* over real-valued smooth functions. A module is like a vector space except over a *ring* (the functions form a ring because there is no multiplicative inverse for any function that has a zero). So vector fields equipped with the commutator form an *algebra over a ring*.

where  $\nabla_X f \equiv X[f] \equiv df(X) = X^\mu \frac{\partial f}{\partial x^\mu}$  for a smooth map  $f : M \rightarrow \mathbb{R}$ . Here, we are considering the chart  $(U, \varphi)$  such that for  $p \in M$ ,  $\varphi(p) = x$ . The components of the Christoffel symbols  $\Gamma_{\mu\nu}^\lambda$  are given by:

$$\nabla_{\partial_\mu} \partial_\nu \equiv \nabla_\mu \partial_\nu = \Gamma_{\mu\nu}^\lambda \partial_\lambda \quad \text{with } \partial_\mu \equiv \frac{\partial}{\partial x^\mu}.$$

For two vector fields  $X = X^\mu \partial_\mu$  and  $Y = Y^\mu \partial_\mu$ , one obtains:

$$(\nabla_X Y)^\nu = X^\mu \partial_\mu Y^\nu + X^\mu \Gamma_{\mu\rho}^\nu Y^\rho.$$

- a. Show that for a vector field  $X = X^\mu \partial_\mu$  the action of the covariant derivative  $\nabla$  on a 1-form  $\omega = \omega_\nu dx^\nu$  is given by: 1pt.

$$(\nabla_X \omega)_\nu = X^\mu \partial_\mu \omega_\nu - X^\mu \Gamma_{\mu\nu}^\lambda \omega_\lambda.$$

- b. Consider another chart  $(V, \psi)$  where  $y = \psi(p)$ . In the region  $V$ , the Christoffel symbols can be written in  $y$ -coordinate components as  $\tilde{\Gamma}_{\alpha\beta}^\gamma$ . Show that in the intersection  $U \cap V$ , this symbol can be written in terms of  $\Gamma$  as:

$$\tilde{\Gamma}_{\alpha\beta}^\gamma = \frac{\partial x^\lambda}{\partial y^\alpha} \frac{\partial x^\mu}{\partial y^\beta} \frac{\partial y^\gamma}{\partial x^\nu} \Gamma_{\lambda\mu}^\nu + \frac{\partial^2 x^\nu}{\partial y^\alpha \partial y^\beta} \frac{\partial y^\gamma}{\partial x^\nu}.$$

Observe that  $\Gamma$  does not transform as a tensor due to the second derivative term. 2pt.

- c. Show that the two-index object,

$$(\nabla_\mu \omega)_\nu = \partial_\mu \omega_\nu - \Gamma_{\mu\nu}^\lambda \omega_\lambda,$$

transforms as components of a tensor. Here,  $\omega = \omega_\nu dx^\nu$  is a 1-form. 1pt.

**Problem 3 Geodesics on  $S^2$ :** The geodesic differential equation has the form:

$$\frac{d^2 x^\mu}{d\tau^2} + \Gamma_{\rho\sigma}^\mu \frac{dx^\rho}{d\tau} \frac{dx^\sigma}{d\tau} = 0,$$

Here, we consider  $S^2$  with the metric  $ds^2 = a^2 d\theta^2 + a^2 \sin^2 \theta d\varphi^2$ , for polar angle  $\theta$ , azimuthal angle  $\varphi$ , and constant  $a > 0$ .

- a. Calculate the Christoffel symbols to show that the above equation takes the following form on  $S^2$ : 2pt.

$$\frac{d^2 \theta}{d\tau^2} - \sin \theta \cos \theta \left( \frac{d\varphi}{d\tau} \right)^2 = 0, \quad \frac{d^2 \varphi}{d\tau^2} + 2 \cot \theta \frac{d\theta}{d\tau} \frac{d\varphi}{d\tau} = 0.$$

- b. Now show that if  $\theta = \theta(\varphi)$  is the equation of a geodesic, then the two equations can be combined into a single equation: 1pt.

$$\frac{d^2\theta}{d\varphi^2} - 2 \cot \theta \left( \frac{d\theta}{d\varphi} \right)^2 - \sin \theta \cos \theta = 0.$$

- c. Now define  $f(\theta) = \cot \theta$ . Show that  $f$  satisfies the differential equation:

$$\frac{d^2 f}{d\varphi^2} + f = 0 .$$

Give the general solution. What do the geodesics on  $S^2$  look like? 2pt.

*Hint: it may help to consider  $f$  in the standard embedding of the sphere in  $\mathbb{R}^3$ .*

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**Problem 4 Curvature of the sphere:**

The Riemann tensor is a (1,3)-tensor given by

$$R(X, Y)Z = \nabla_X \nabla_Y Z - \nabla_Y \nabla_X Z - \nabla_{[X, Y]} Z ,$$

for  $X, Y, Z \in TM$ .

- a. Write the components of the Riemann tensor

$$R^k{}_{\ell ij} \equiv dx^k (R(\partial_i, \partial_j)\partial_\ell)$$

in terms of partial derivatives and Christoffel symbols. 1.5pt.

- b. Show that the Riemann tensor vanishes for flat Euclidean space  $ds^2 = dx^2 + dy^2$ . 1pt.

- c. Calculate the Riemann tensor for the sphere  $S^2$ , where  $ds^2 = a^2 d\theta^2 + a^2 \sin^2 \theta d\varphi^2$ . What is the  $a$ -dependence for  $R^k{}_{\ell ij}$ ? The Riemann tensor with only lower indices is defined as:  $R_{klij} = g_{mk} R^m{}_{\ell ij}$ . What is the  $a$ -dependence of  $R_{klij}$ ? 2.5pt.

*Hint: there should only be one independent, non-zero term after considering symmetries in the components of the Riemann tensor.*

- d. Calculate the Ricci tensor  $R_{ij} = R^k{}_{ikj}$  and the Ricci scalar  $R = g^{ij} R_{ij}$ . What is the  $a$ -dependence of the Ricci scalar and how does this scalar vary over the sphere? 2pt.